Modelling high-resolution spatially-resolved SNR spectra with the Sardinia Radio Telescope



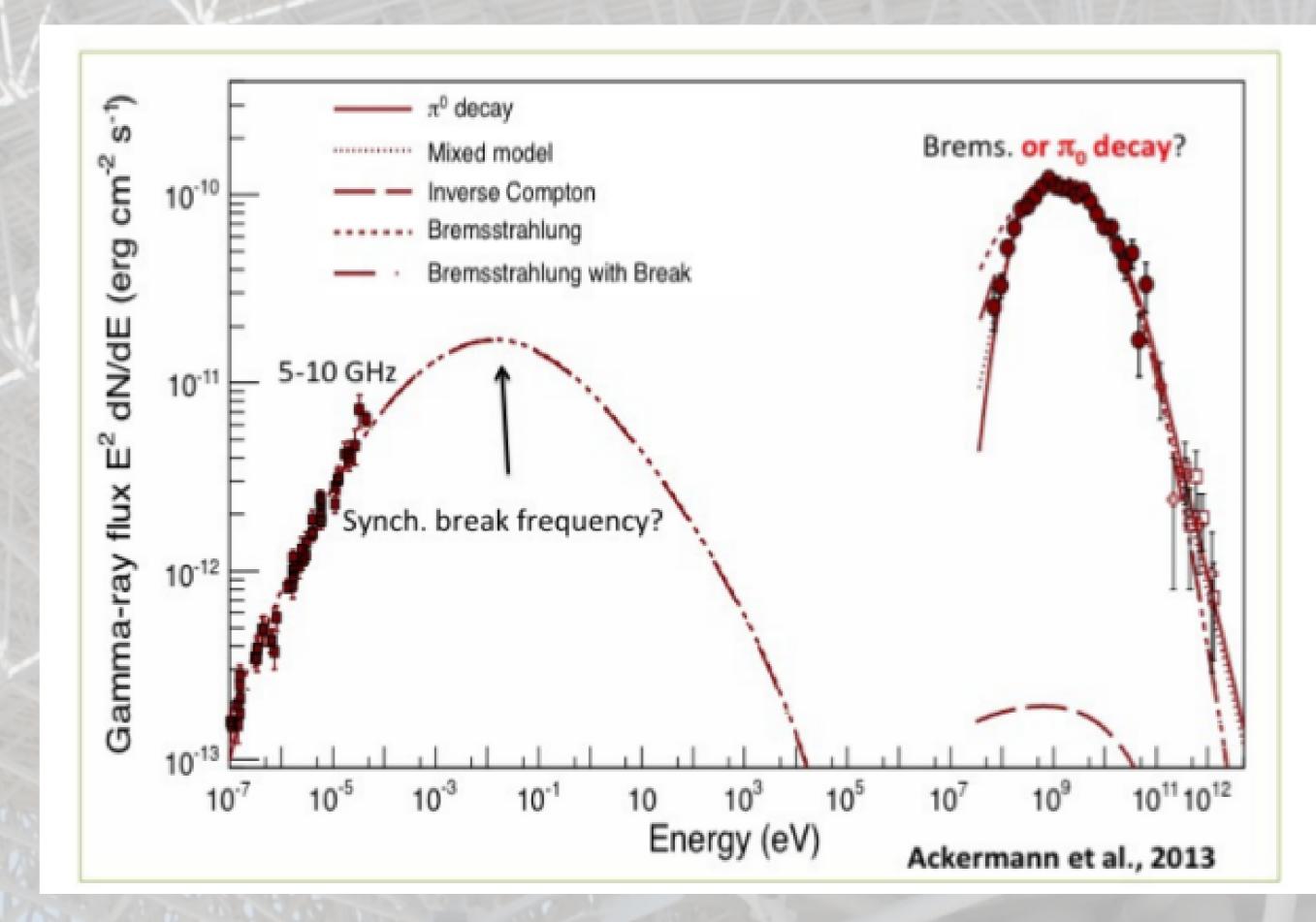
Sara Loru $^{(1)}$ (3), Alberto Pellizzoni $^{(1)}$, Elise Egron $^{(1)}$, Noemi Iacolina $^{(1)}$, Simona Righini $^{(2)}$, Marco Marongiu $^{(1)}$, Sara Mulas $^{(3)}$, Giulia Murtas $^{(3)}$, Davide Simeone $^{(3)}$, Maura Pilia $^{(1)}$, Matteo Bachetti $^{(1)}$, Alessio Trois $^{(1)}$, Roberto Ricci $^{(2)}$, Andrea Melis $^{(1)}$, Raimondo Concu $^{(1)}$ ⁽¹⁾ INAF-Observatory of Cagliari, Italy ⁽²⁾ INAF- Istituto di Radioastronomia, Bologna, Italy Abstract

Supernova Remnants exhibit spectra typically featured by synchrotron emission mostly arising from the relativistic emitting electrons in the radio band and high-energy emission from both leptonic processes like Bremsstrahlung and Compton Inverse produced by the radio-electrons interacting with ambient photons, and hadronic process of π^0 mesons decay which is a direct signature of Cosmic Rays accelerated in Supernova Remnants. Thanks to radio imaging observations obtained in three frequency bands (1.4, 7, 22 GHz) with the Sardinia Radio Telescope (www.srt.inaf.it), we can model separately different SNR regions. Indeed, in order to disentangle interesting and peculiar hadron contributions in the high-energy spectra and better constrain SNRs as Cosmic Rays emitters, it is crucial to fully constrain lepton contributions first through radio-observed parameters. In particular, the Bremsstrahlung and Inverse Compton bumps observed in gamma-rays are typically bounded to synchrotron spectral slope and cut-off in the radio domain, associated with different SNR regions and electron populations.

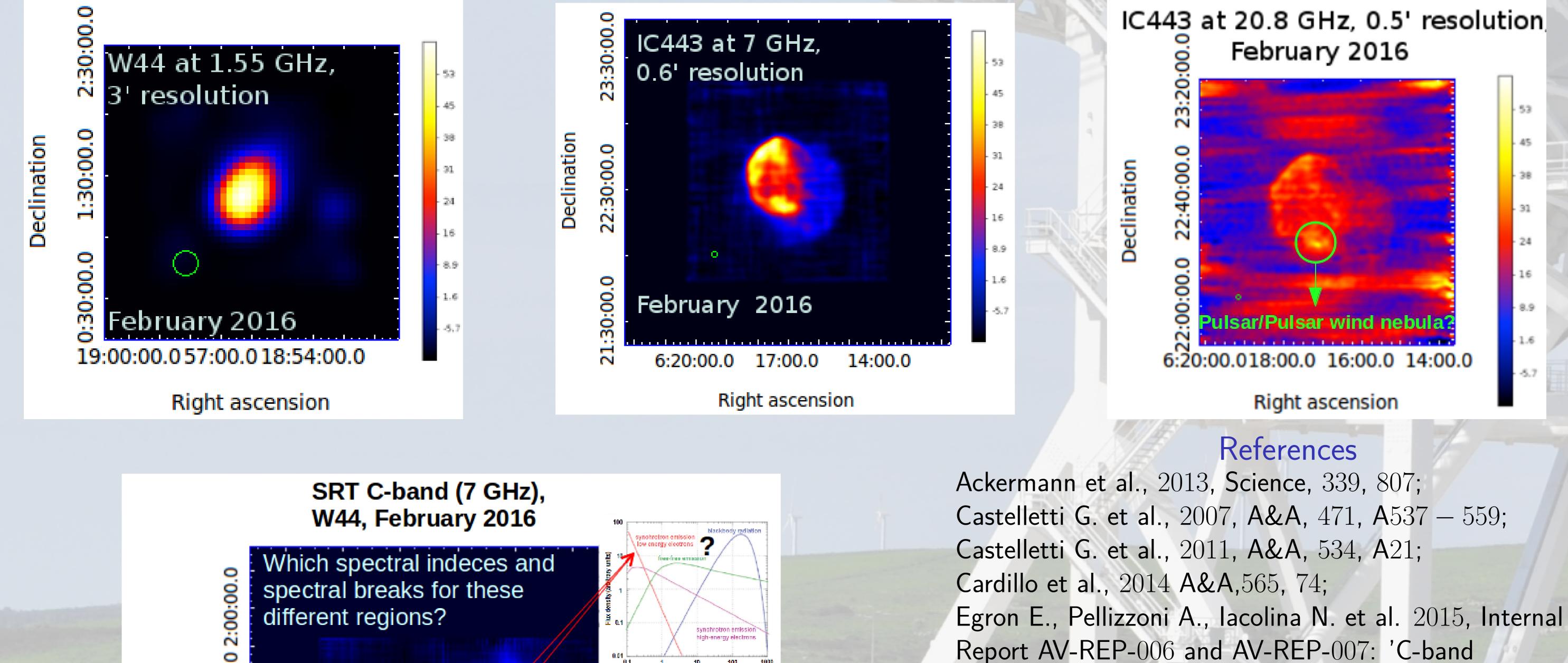
• Synchrotron-emitting electrons in the high-frequency radio band are

SR'

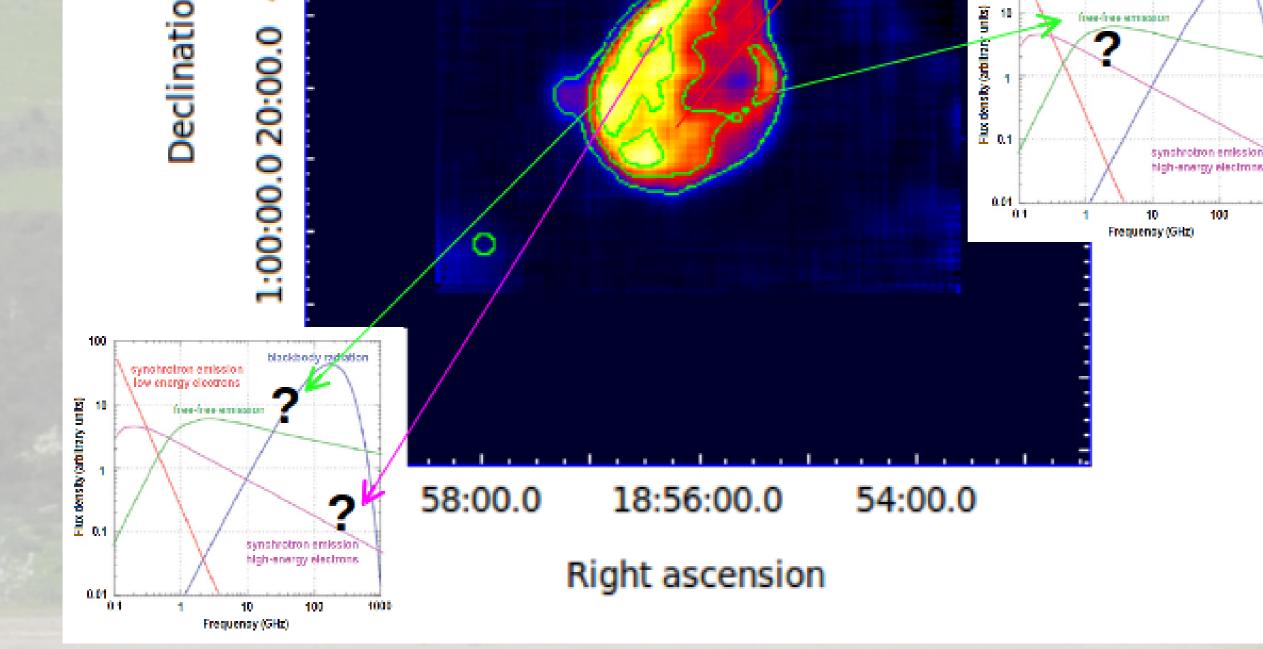
- responsible for the observed high-energy phenomenology as -e.g.- Inverse Compton and Bremsstrahlung emission components observed in gamma-rays, to be disentangled from hadrons emission contribution (providing constraints on the origin of cosmic rays).
- Multiwavelength data on SNRs are sparse and no spatially-resolved spectra are available in the 5 - 20 GHz range, critical for model assessment, also for the most studied and bright objects
- Recent constraints on Cosmic Rays emission from SNRs and related models are based on integrated radio fluxes only implying the simplistic "single-zone" assumption of a single electron population for the whole SNR.
- With SRT we perform accurate on-the-fly scans of W44 and IC443 SNRs at three frequencies (L, C, K bands) in order to obtain detailed radio images and spatial-resolved spectral-slope measurements (synchtron spectral breaks are possibly expected in this range).



SRT multifrequency maps of W44 and IC443



Ξ



Handbook of Frequency Allocations and Spectrum Protection for Scientific Uses: Second Edition; Reynolds S.P., 2008, Ann. Rev. Astron. Astrophys., 46, 89 - 126;Gao X.Y., Han J.L., Reich W. et al. 2011, A&A, 5291, A159; Giuliani et al., 2011, ApJ, 742, L30; Green D.A. 2009, Bull. Astron. Soc. India, 37, 45; Green D.A. 2014, A Catalogue of Galactic Supernova Remnants (2014 May version) (http://www.mrao.cam.ac.uk/surveys/snrs/); Sun X. H., Reich P., Reich W. et al. 2011, A&A, 536, A83; Elise Egron: egron@oa-cagliari.inaf.it

Observations of Supernova Remnants with SRT';

Sara Loru: saraloru@oa-cagliari.inaf.it

Alberto Pellizzoni:

low energy electron

apellizz@oa-cagliari.inaf.it